

1 The Great Balls of Fire: A probabilistic approach to
2 quantify the hazard related to ballistics — a case study
3 at La Fossa volcano, Vulcano Island, Italy

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13 **Abstract**

14 Volcanic ballistic projectiles (VBPs) represent a common hazard around vol-
15 canoes associated with explosive eruptive styles, and although the number of
16 impact on the ground decreases exponentially with distance from the vent,
17 exposed elements typically present a high vulnerability to VBPs. Recent ef-
18 forts to assess the hazard related to VBPs have mostly been deterministic,
19 with the final aim being hazard zoning. This approach, suitable when no
20 human settlement is present within a radius of a few kilometres around the
21 vent, is of limited use when urban areas extend close to the active volcanic
22 centre, which requires probabilistic hazard assessments in order to plan and
23 prioritize emergency actions. This is the case at La Fossa volcano, the ac-
24 tive crater of the island of Vulcano (Italy), where the activity of the past
25 1,000 years is characterized by intense ejection of VBPs and where human
26 settlements are located as close as 750 m from the vent.

27 To address the need of a probabilistic hazard assessment for VBP im-
28 pacts, we developed a new numerical model named *The Great Balls of Fire*
29 (GBF), which describes the ballistic trajectory of particles with variable drag
30 coefficients in a stochastic way and uses a digital elevation model (DEM) to
31 account for topographic barriers. In parallel, a set of associated functions
32 post-processes the output of the model to produce a comprehensive prob-

33 abilistic hazard assessment for VBP impacts. Outcomes include probabil-
34 ity maps to exceed given thresholds of kinetic energies at impact, hazard
35 curves and probabilistic isoenergy maps. Probabilities are calculated either
36 on equally-sized pixels or zones of interest.

At La Fossa, the last 1,000 years of activity were characterized by at least 8 long-lasting Vulcanian cycles, each associated with the ejection of VBPs. The GBF model was validated using VBPs produced during the last 1888–1890 eruption that had sufficient stratigraphic constraints to discard reworking, and a generic probabilistic eruption scenario was identified for future Vulcanian eruptions. Results suggest a $10^{-3} - 10^{-2}\%$ probability of occurrence of VBP impacts with kinetic energies larger than 4,000 J at the touristic locality of Porto. Additionally, the physical vulnerability of the built environment was assessed from field observations and combined with published literature to identify typical fragility curves for roof perforation. Hazard and vulnerability assessments were then combined to provide a first estimate of the potential impact of VBPs during future Vulcanian eruptions, indicating that for a 50% probability of occurrence of the hazard, half of the building stock has a $10^{-4} - 10^{-3}\%$ probability of roof perforation.

37 *Keywords:* Probabilistic hazard assessment, Volcanic ballistic projectiles,
38 Pre-event impact assessment, Physical vulnerability, Roof penetration,
39 Vulcano Island, La Fossa

40 1. Introduction

41 Volcanic ballistic projectiles (VBP) are angular to sub-angular fragments
42 that decouple from the jet phase of explosive events to follow a near-ballistic
43 trajectory modified by drag forces (Alatorre-Ibargüengoitia et al., 2012).
44 VBPs can be distinguished between blocks, typically of angular shape and
45 lithic origin, and bombs, typically of rounded shape and juvenile origin.
46 These ballistic projectiles can be produced in all types of volcanic eruptions,
47 but are particularly abundant with Vulcanian and Strombolian styles (e.g.
48 Vanderkluysen et al., 2012). Because of their high kinematic energies and
49 temperatures, VBPs constitute a major threat in proximal areas due to their
50 high potential impact to life and the built environment and their propensities
51 to ignite fires. As examples, Pomonis et al. (1999) reports that bombs lighter
52 than 1 kg are known to have penetrated thatched and galvanized iron roofs
53 during previous eruptions of Furnas volcano (Azores), and Pistolesi et al.

54 (2011) and Rosi et al. (2013) reported wildfires triggered by incandescent
55 blocks during the 2007 crisis of Stromboli.

56 Numerous models for ballistic ejection have been developed since the
57 1940's to primarily invert field observations and estimate eruptive conditions
58 (e.g. ejection velocity, i.e. Fudali and Melson, 1971; Wilson, 1972; Minakami,
59 1942; Steinberg and Lorenz, 1983). Although accounting for drag effects,
60 these first models considered the ejection of blocks into a still atmosphere,
61 commonly leading to an overestimation of the atmospheric drag forces and,
62 consequently, unrealistically high ejection velocities. In the context of Vulca-
63 nian eruptions, later models considered the presence of a caprock accelerated
64 by the expansion of underlying gas behaving as a coherent plug until a max-
65 imum velocity is reached, at which point the fractured caprock disaggregates
66 and individual ballistic blocks are released (Self et al., 1979; Wilson, 1980;
67 Fagents and Wilson, 1993). This disaggregation height has been recently
68 suggested to occur when the acceleration is 8% of the initial acceleration of
69 the caprock (Alatorre-Ibargüengoitia et al., 2012). This implies a region of
70 reduced drag in the vicinity of the eruptive source, within which the air will
71 itself be moving radially from the source at a velocity comparable to that
72 of the clasts (Fagents and Wilson, 1993). Using this concept, the effect of
73 drag becomes important only when the velocity of the clast gradually decou-
74 ples from that of the surrounding air, which allows to reproduce observed
75 deposits with significantly lower ejection velocities. The *Eject!* model of
76 Mastin (2001) is the most widely used model for hazard zonation, and allows
77 to assess the maximum distance that a block of a given size and density can
78 reach for given sets of eruptive conditions (e.g. ejection speed and angle), and
79 accounts for a region of reduced drag defined as an arbitrary distance above
80 the vent and a variation of drag from ejection to landing. De' Michieli Vitturi
81 et al. (2010) proposed a coupled Eulerian-Lagrangian model to describe the
82 dynamics of large particles during Vulcanian eruptions, providing a detailed
83 parametrization of the complex radial and vertical acceleration and deceler-
84 ation patterns of the initial jet phase. Alatorre-Ibargüengoitia et al. (2012)
85 presented a model coupling lab measurement of the effect of shape on the
86 drag of volcanic particles and a caprock model relating the energy consump-
87 tion required by fragmentation to the ejection velocity of ballistics (Alatorre-
88 Ibargüengoitia and Delgado-Granados, 2006; Alatorre-Ibargüengoitia et al.,
89 2010). Recently, Tsunematsu et al. (2014) developed a new approach ac-
90 counting for multiple particles and collision between blocks.

91 The aim of hazard assessments is to quantify the probability of occurrence

92 of a hazardous phenomenon of a given magnitude (Fournier d’Albe, 1979;
93 Mendoza-Rosas and la Cruz-Reyna, 2008). In volcanology, where eruptions
94 constitute a multi-hazard system, this process is commonly achieved by i)
95 the field characterization of the deposits in order to constrain and quantify
96 Eruption Source Parameters (ESPs), ii) the compilation of a catalogue of
97 eruptions and phenomena at a given volcano to infer eruption scenarios and
98 iii) the forward modelling of a given phenomenon using appropriate models
99 (e.g. Biass et al., 2014). Recent hazard assessments in all fields of natural
100 hazards increasingly rely upon probabilistic techniques in order to account
101 for the inherent uncertainty of natural processes (e.g. Geist and Parsons,
102 2006; Gonzalez et al., 2009; Heneka and Hofherr, 2011). In volcanology,
103 stochastic strategies have been widely applied to the modelling of tephra (e.g.
104 Bonadonna, 2006; Jenkins et al., 2012) and, more recently, lava flows (e.g.
105 Connor et al., 2012). Probabilistic hazard assessments typically require the
106 identification of eruption scenarios, characterized by typical ranges of ESPs.
107 Each ESP is defined as a probability distribution, constrained by minimum
108 and maximum values as well as a distribution shape. Hazard assessments for
109 ballistics are, however, often based on a deterministic definition of eruption
110 scenarios aiming at producing hazard zones for different block size, ejection
111 angle and initial velocities (Sandri et al., 2014; Alatorre-Ibargüengoitia et al.,
112 2006, 2012). Recently, Fitzgerald et al. (2014) proposed a new probabilistic
113 approach based on the model of Tsunematsu et al. (2014), in which crucial
114 ESPs were quantified in terms of mean value and standard deviation from
115 the field study of 3587 impact craters.

116 We propose here a new approach to assess the hazard and the impact
117 on the built environment related to the ejection of ballistic blocks, compiled
118 as a package called *Great Balls of Fire* (GBF; Lewis & Hammer, 1957, Sun
119 Studio). The first part of the GBF package comprises a model written in
120 Scala, with the main features being i) the stochastic sampling of ESPs, ii)
121 the implementation of a variable drag coefficient, iii) the ability to use a DEM
122 to account for topographic barriers and iv) the possibility to work on a single
123 CPU or on a cluster of computers. The second part of the package provides
124 *Matlab* routines to post-process the output of the model in order to compute
125 the geographical probability to be affected by VBP impacts of a given en-
126 ergy threshold, and allows to easily export the results into GIS platforms.
127 This paper first describes the ballistic model, which is then tested and vali-
128 dated using field measurements of VBPs produced during the last Vulcanian
129 eruption of La Fossa Volcano, Vulcano Island, Italy. We then compiled an

130 eruption scenario for a Vulcanian activity and applied the method to compile
131 probabilistic hazard maps for the ejection of ballistics at La Fossa. Outcomes
132 are combined with a rapid assessment of the built environment to produce a
133 first-order pre-event impact assessment of the buildings stock.

134 **2. Case study of Vulcano Island**

135 Vulcano is the southernmost islands of the Aeolian archipelago and, along
136 with Lipari and Stromboli, one of the active volcanic systems of the archipelago
137 (De Astis et al., 1997; Gioncada et al., 2003, Fig. 1). The sub-aerial activity
138 of Vulcano started between 135 and 120 ka years ago (Zanella et al., 2001),
139 after which volcanism migrated towards N–NW through time, generating a
140 composite structure characterized by four, juxtaposed volcanic edifices in-
141 cluding the cone of La Fossa, center of the current activity (Keller, 1980;
142 Frazzetta et al., 1983; De Astis et al., 1997, 2013; Gioncada et al., 2003).
143 The eruptive history and structure of the 391 *m*-high La Fossa cone was
144 mostly studied by Arrighi et al. (2006), De Astis et al. (2013), Dellino et al.
145 (2011), Di Traglia et al. (2013), Frazzetta et al. (1984), Frazzetta et al. (1983)
146 and Keller (1980).

147 The eruptive history of the last 1,000 years was reconstructed based
148 on stratigraphic studies (De Astis et al., 2013; Di Traglia, 2011) and his-
149 torical chronicles (De Fiore, 1922; Mercalli and Silvestri, 1891). Following
150 the nomenclature of Di Traglia et al. (2013), the most recent deposits were
151 grouped in two stratigraphic clusters including the Palizzi–Commenda Eruptive
152 Cluster (PCEC) and the Gran Cratere Eruptive Cluster (GCEC).

153 The PCEC is itself divided in two sub-units, including the Palizzi and
154 the Commenda units. The Palizzi unit is characterized by both explosive and
155 effusive styles interrupted by periods of major erosion and re-sedimentation
156 of the newly deposited material (Frazzetta et al., 1984, 1983; Di Traglia,
157 2011; Dellino and La Volpe, 1997; Dellino et al., 2011; De Astis et al., 2013).
158 The Palizzi unit started with a long-lasting emission of mafic ash followed by
159 a first rhyolitic sub-Plinian eruption (PAL B; Di Traglia et al., 2013). After
160 a period of quiescence occurred a second long-lasting activity of ash emission
161 and a new sub-Plinian phase of trachitic composition (PAL D; Di Traglia
162 et al., 2013), followed by the effusion of two lava flows. The Commenda unit
163 is a magmatic–hydrothermal eruption (Gurioli et al., 2012) comprising the
164 Breccia di Commenda deposit (~1240AD), characterized by fallouts with a
165 high lithic to juvenile ratio and anomalously large amounts of S and metals

166 and the ejection of dense lithic VBPs (Di Traglia et al., 2013; Gurioli et al.,
167 2012).

168 The GCEC (1440AD–1890AD; Di Traglia et al., 2013) started with a
169 steam–blast eruption on the 5th of February 1444 (Mercalli and Silvestri,
170 1891), and around 1550AD occurred the first of the eight Vulcanian erup-
171 tions of the GCEC (Di Traglia et al., 2013). The last eruption occurred in
172 1888–1890 and was characterized by plume heights between 1 and 10 *km*. An
173 intense ejection of VBPs was reported throughout the ~ 2 year–long cycle,
174 which produced different morphologies throughout the eruption, with dense
175 lithic blocks occurring at the beginning and the end of the cycle and juve-
176 nile bread–crust bombs ejected mostly at the half of the eruption (Bianchi,
177 2007; Di Traglia, 2011). Outcrops with VBPs associated with the 1888–1890
178 eruption are shown in Figure 1 (S1–S3). In addition, historical reports also
179 mention that a warehouse located close to the so-called Stevenson Castle
180 (pink star on Fig. 1) was impacted by a VBP.

181 About 800 people permanently live on Vulcano, but daily peaks can reach
182 20’000 during the summer season. Four settlements are present on the island.
183 In the south, Piano lies on top of the filled caldera of Vulcano Primordiale and
184 is the home of most of the permanent inhabitant. The remaining settlements
185 of the Porto area, Vulcanello and Lentia comprise most of the hotels and
186 tourism facilities. The topography (Fig. 1) suggests that Piano and Lentia
187 are sheltered by barriers, whereas the Porto and Vulcanello areas lie on a
188 plain directly North of the La Fossa cone.

189 3. The GBF Model

190 The GBF model is based on classical movement equations using gravity
191 and drag force. The model, described below, accounts for a standard at-
192 mosphere, the influence of the wind and a region of reduced drag following
193 Mastin (2001). The bomb simulator was implemented using the Scala lan-
194 guage and parallelized with the Akka actor framework. User interactions
195 are provided through a minimalist command line interface and all simulation
196 settings are defined in a simple and human readable configuration file.

197 3.1. Governing equations

198 Each particle is approximated by a sphere and described by a mass m ,
199 an average diameter D , a position \mathbf{r} and a velocity \mathbf{v} . The VBP trajectory
200 is described by the following equations:

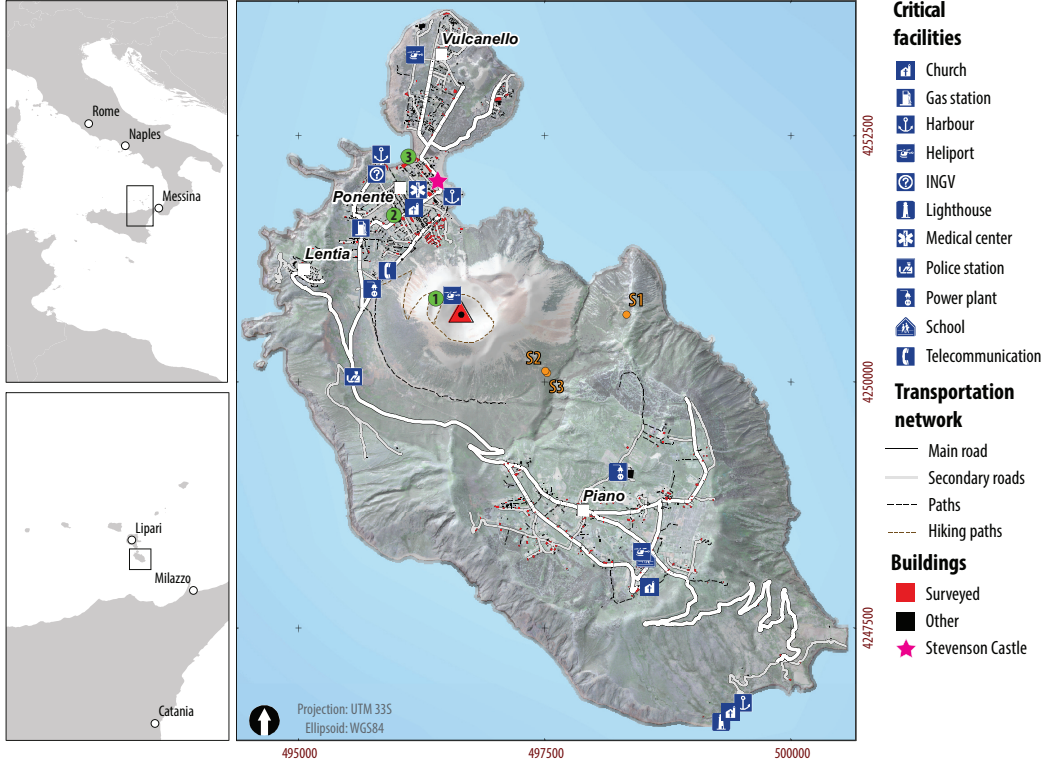


Figure 1: Overview of Vulcano Island, showing the road network, the location of critical infrastructures and the buildings footprints. Green dots show the reference points used for the sensitivity analysis. Orange dots show the field location of the sampling sites. Adapted from Biass et al. (submitted).

201
$$\mathbf{u} = \mathbf{v} - \mathbf{w} \tag{1}$$

$$\ddot{\mathbf{r}} = \dot{\mathbf{v}} = \mathbf{a} = \frac{-\rho_a A C_d \mathbf{u} |\mathbf{u}|}{2m} + \mathbf{g} \tag{2}$$

202 where A is the fluid cross area, C_d the drag coefficient, ρ_a the air density,
 203 \mathbf{u} the velocity of the VBP relatively to the wind \mathbf{w} and \mathbf{g} the acceleration
 204 gravity vector. The computation of the drag coefficient and the air density
 205 depends on the bomb altitude and velocity. For a given altitude z , the air
 206 temperature T and pressure p are computed using the following formulas:

$$T(z) = T_0 + \gamma z \tag{3}$$

$$p(z) = p_0 \left(\frac{T(z)}{T_0} \right)^{-\frac{g}{R\gamma}} \quad (4)$$

207 where T_0 and p_0 are respectively the air temperature and pressure at
 208 sea level, γ is the thermal lapse and R the gas constant. This allows the
 209 computation of both the air density and the kinematic viscosity ν_a :

$$\rho_a(z) = \frac{p(z)}{RT(z)} \quad (5)$$

$$\nu_a(z) = \left(\frac{6.70810^{-3}}{T(z) + 117} \right) \cdot \left(\frac{T(z)}{273} \right)^{\frac{3}{2}} \quad (6)$$

210 The particle Reynolds number, based on the air characteristics detailed
 211 above and the bomb diameter and speed, is used to determine the drag
 212 coefficient C_d :

$$Re = \frac{\rho_a u D}{\nu_a} \quad (7)$$

$$C_d = \begin{cases} 0.1 & \text{if } Re < 3 \times 10^5, \\ 0.5 & \text{else.} \end{cases} \quad (8)$$

213 Since bombs are ejected together with an expanding mass of gas, the drag
 214 coefficient may be reduced according to the following equation:

$$C'_d = \begin{cases} C_d \left(\frac{r}{r_d} \right)^2 & \text{if } r < r_d, \\ C_d & \text{else.} \end{cases} \quad (9)$$

215 3.2. Random bomb generation

216 In order to produce probabilistic hazard assessments for VBPs based on
 217 identified scenarios, the GBF model is implemented with a module for gen-
 218 erating bombs with random initial conditions. Each VBP is generated with
 219 ESPs sampled stochastically and constrained either on Gaussian or uniform
 220 distributions (Table 1). Each VBP is characterized by a diameter and a
 221 mass, which, assuming a spherical shape, are used to calculate the mass.
 222 Additional tests are performed to ensure that all constraints in Table 1 are
 223 satisfied, else all parameters are discarded and re-sampled.

Table 1: Summary of parameters modelled stochastically in the GBF model. $N(\mu, \sigma)$ represents a normal distribution with average μ and standard deviation σ . $U(a, b)$ represents an uniform distribution with values in the interval $[a, b]$.

Parameter	Distribution	Constraint
Ejection velocity (v)	$v \in N(v_\mu, v_\sigma)$	$v > 0$
Ejection angle (ϕ)	$\in N(\phi_\mu, \phi_\sigma) $	
Ejection azimuth (θ)	$\in U(0, 2\pi)$	
Density (d)	$\in N(d_\mu, d_\sigma)$	$d > 0$
Grain size (Φ)	$\in N(\Phi_\mu, \Phi_\sigma)$	

224 3.3. Numerical model and implementation

225 Equations 1–9 are solved numerically using Runge-Kutta 4th order with
 226 a time step $\Delta t = 0.01$ s. In the absence of an analytic solution, we tested the
 227 accuracy of the output by solving the trajectories of 10 000 randomly sampled
 228 VBPs with time steps of 0.01 s and 0.001 s. Using the smaller time step as
 229 a reference, we computed the absolute error as the distance between impact
 230 points under both conditions. The error was below 1 m for 99.56 % of the
 231 bombs and the maximum recorded error was below 3 m . When normalized
 232 by the distance between the impact and the vent, only 9 bombs out of 10,000
 233 had a relative error of more than 0.01 %.

234 3.4. Validation with field data

235 The GBF model was validated using the field observation of six VBPs
 236 associated with the 1888–1890 eruption presenting sufficient stratigraphic
 237 constraints to discard possible reworking and displacement. The VBPs were
 238 classified in three typical morphologies including i) lithic blocks, either fresh
 239 or altered , ii) thin rind breadcrust bombs and iii) thick rind breadcrust
 240 bombs.

241 Firstly, the S1 sampling site (Fig. 1; Table 2) represents the thick-rinded
 242 breadcrust bomb identified by Bianchi (2007) characterized by a diameter
 243 of 25 cm and a density of 1800 $kg\ m^{-3}$ located ~ 1560 m from the vent.
 244 Using the *Eject!* model to infer initial conditions required to reproduce the
 245 observation, Bianchi (2007) identified two end-member scenarios. On one
 246 hand, the minimum required initial velocity was assessed using an ejection
 247 angle of 45° from vertical, leading to an ejection velocity of 145 $m\ s^{-1}$. Based
 248 on the observations of Mercalli and Silvestri (1891) on the crater morphology
 249 during the 1888–1890 eruption, who describe the crater as surrounded by
 250 very steep slopes, an inclination of 15° from the vertical was also tested to

251 represent a more realistic ejection angle, which results in an initial velocity
 252 of 350 m s^{-1} . Such a velocity is higher than the typical range reported in
 253 the literature for Vulcanian explosions (e.g. Alatorre-Ibargüengoitia et al.,
 254 2012). Secondly, the S2 sampling site represents a $20 \times 20 \text{ m}$ area where
 255 the populations of different VBPs morphologies were studied. From a total
 256 of 111 VBPs found in the area, the S2 sampling site shows a dominance of
 257 lithic blocks (80%) with minor thin- (14%) and thick-rinded (6%) breadcrust
 258 bombs. The most representative VBP of each morphology was measured in
 259 terms of the length of the 3 main orthogonal axes and an equivalent diameter
 260 was calculated using the geometric mean (Table 2). Finally, one abnormally
 261 large thick-rinded breadcrust bomb was located 40 m further (sampling site
 262 S3) and characterized in a similar way (Table 2).

263 We used the GBF model to invert these observations to infer the po-
 264 tential range of initial conditions. Sets of simulations of 10^5 particles were
 265 performed varying the initial velocity between 100 and 350 m s^{-1} (increment
 266 of 25 m s^{-1}) and the ejection angle between 5 and 45° from the vertical (in-
 267 crement of 5°). The density was also adapted for each VBP morphology and
 268 was randomly sampled for each simulated bomb on a Gaussian distribution
 269 (Table 2). At each simulation, both initial velocity and ejection angle were
 270 allowed a narrow variation characterized by a standard deviation equal to
 271 half of the increment. From the results, the mean distance reached by VBPs
 272 was calculated for each sets of initial velocity and ejection angle. Figure 2
 273 represents the difference between the distances of the observed VBP and the
 274 mean distance obtained by the GBF model as a function of initial velocity
 275 and ejection angle. The minimum discrepancy between observed and com-
 276 puted VBPs is given by the 0 line in Figure 2 for radii of reduced drags of
 277 200 m (black line), 600 m (blue line) and 1000 m (red line).

278 For the S1 sample, both the GBF and *Eject!* models result in similar
 279 minimum conditions, i.e. a velocity of 145 m s^{-1} for an ejection angle of 45°
 280 (Fig. 2). In contrast, the GBF model suggests a velocity of $\sim 225 \text{ m s}^{-1}$ for
 281 an angle of 15° , which is significantly lower than the 350 m s^{-1} suggested by
 282 Bianchi (2007) but more realistic when compared to typical ejection veloci-
 283 ties reported for Vulcanian explosions (Alatorre-Ibargüengoitia et al., 2006,
 284 2012). Nevertheless, due to the location of the S1 sample (i.e. on the edge of
 285 the Piano caldera, 1.6 km away from the vent) and the absence of historical
 286 report of VBP reaching the Piano caldera, we assume the S1 sample as an
 287 extreme case-figure. The S2 and S3 samples are well reproduced by the GBF
 288 model (Fig. 2), where a sound exit velocity of 150 m s^{-1} typically results in

Table 2: Summary of observed VBPs associated with the 1888–1890 eruption used for the field validation of the GBF model. The distance represents the euclidean distance from the actual vent. The sample locations are reported on Fig. 1. *BCB* stands for breadcrust bomb.

Sampling site	Type	Distance (m)	Axes lengths (cm)	Diameter (cm)	Density (kg m ⁻³)	
					Mean	σ
S1	Thick-rinded BCB	1560	—	25	1600	200
S2	Altered block	960	120 × 65 × 40	68 ^a	2500	100
S2	Fresh block	960	40 × 35 × 22	31 ^a	2500	100
S2	Thin-rinded BCB	960	47 × 30 × 10	24 ^a	800	50
S2	Thick-rinded BCB	960	35 × 30 × 18	27 ^a	1600	200
S3	Thick-rinded BCB	1000	70 × 50 × 50	56 ^a	1600	200

^a: Equivalent diameter expressed as the geometric mean of the three orthogonal axes.

289 ejection angles lower than 15–20°. Two extra points can be observed from
 290 Figure 2. Firstly, for a similar equivalent diameter and the same sampling
 291 point, thin-rinded breadcrust bombs require higher ejection velocities than
 292 thick-rinded breadcrust bombs to reproduce the observations, which is due
 293 to the lower kinetic energy of lighter bombs. Secondly, an increased radius
 294 of reduced drag has an overall low effect on the model results, although the
 295 effect increases when reproducing impacts at larger distances from the vent
 296 (e.g. S1) or for lighter VBPs (thin-rinded breadcrust bomb of S2).

297 4. Application to La Fossa volcano

298 4.1. Eruptive scenarios

299 During the activity of the last 1,000 years at La Fossa, two main eruptive
 300 styles were recognized to produce VBPs, including non-juvenile steam blast
 301 eruptions (i.e. Commenda unit) and Vulcanian eruptions (Di Traglia et al.,
 302 2013; De Astis et al., 2013). Since the eruptive dynamics behind both styles
 303 are different, we only consider a Vulcanian-type scenario here.

304 In probabilistic hazard assessments, ESPs are typically characterized by
 305 distributions of values rather than single values fixed deterministically in
 306 order to account for various types of uncertainties (e.g. aleatoric and epis-
 307 temic). Here, we developed a Vulcanian-type scenario built around the ref-
 308 erence 1888–1890 eruption. ESPs were constrained based on the detailed
 309 study of the blocks and bombs provided by the works of Bianchi (2007) and
 310 Tsunematsu (2012) as well as additional fieldwork. During the 1888–1890
 311 eruption, the three types of VBP occurred in different proportions and at
 312 different stages of the eruption. Dense juvenile blocks represent 70–90% of
 313 the observed VBPs on the field and are reported to have occurred mainly at

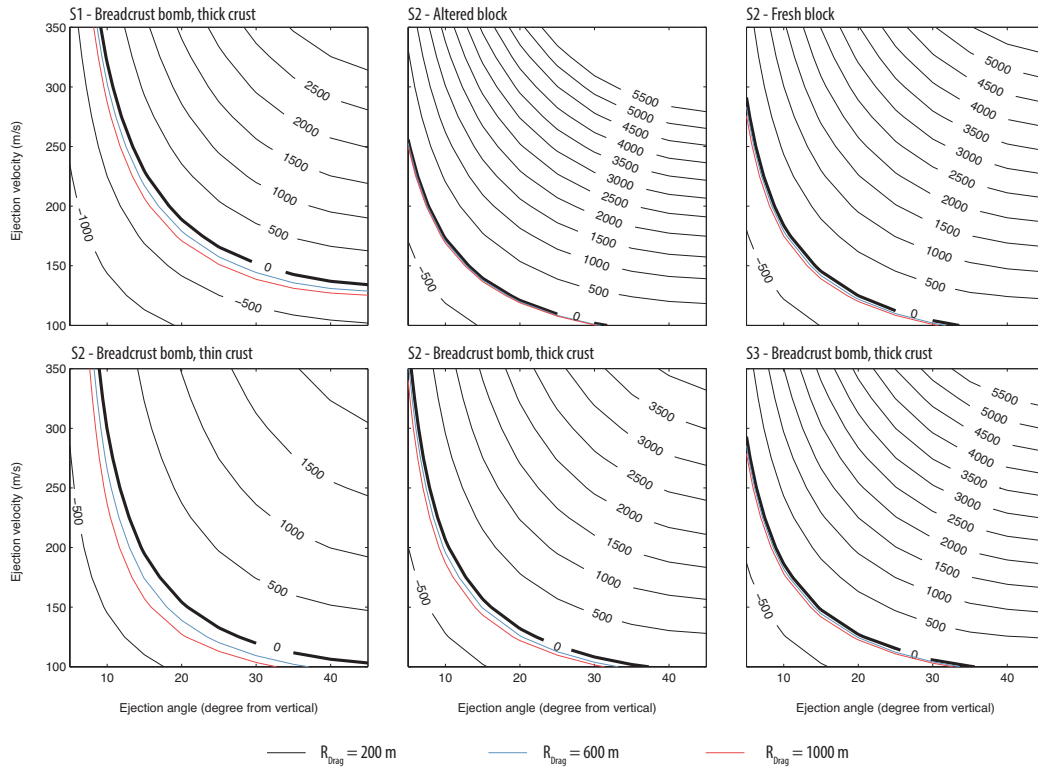


Figure 2: Difference between the distances of the observed VBPs and the mean distance obtained by the GBF model as a function of initial velocity and ejection angle for the various VBP morphologies observed at sampling sites shown in Fig. 1. Various radii of reduced drag are considered and include 200 m (black lines), 600 m (blue line) and 1,000 m (red line). The differences in distances are expressed in metres.

314 the start and the end of the eruption (Mercalli and Silvestri, 1891; Bianchi,
 315 2007; Di Traglia, 2011). Thin-rinded breadcrust bombs represent 5–15% of
 316 the observed VBPs and were produced during the moderate explosions that
 317 occurred at the half of the eruption (Mercalli and Silvestri, 1891; Bianchi,
 318 2007; Di Traglia, 2011). Finally, thick-rinded breadcrust bombs are the ex-
 319 pression of the most intense explosions that occurred throughout the cycle,
 320 and represent 10–20% of the observed VBPs. Since proportions of each VBP
 321 type obtained at the sampling site S2 (Fig. 1; Sect. 3.4) represents a median
 322 value compared the reports of previous authors (Mercalli and Silvestri, 1891;
 323 Bianchi, 2007; Di Traglia, 2011), we assume a proportion of 80% of lithic
 324 blocks, 14% of thin-rinded and 6% of thick-rinded (6%) breadcrust bombs.

325 Table 3 summarizes the ESPs for the Vulcanian eruption scenario at La
 326 Fossa. Variable parameters include i) density ($kg\ m^{-3}$), ii) VBP diameter
 327 (ϕ), iii) ejection velocity ($m\ s^{-1}$) and iv) ejection angle (i.e. azimuth, $^\circ$ from
 328 vertical). The number of observations being too limited to estimate complex
 329 probability distributions (e.g. based on Tsunematsu, 2012, $n = 12$ for density
 330 measurements and $n = 40$ for diameter measurements), we used Gaussian
 331 distributions centred on the mean value (μ) and expressing the uncertainty
 332 using the standard deviation (σ), which accounts for about 68.3% of the
 333 population.

334 Here, the three types of VBPs are characterized by various ranges of den-
 335 sities, and discontinuities discard the possibility of modelling the density as a
 336 continuous range. Instead, separate runs were performed for each VBP type
 337 in which i) the density range was adjusted and ii) the number of particles
 338 simulated was scaled to reproduce the proportions of each VBP type. The
 339 mean densities and associated standard deviations of blocks, thin-rinded
 340 and thick-rinded breadcrumb bombs were set to 2300 ± 100 , 800 ± 50 and
 341 $1600 \pm 200\ kg\ m^{-3}$, respectively. The diameter is expressed on a Gaussian
 342 distribution in ϕ units, which results in a log-normal distribution when con-
 343 verted to metres. The mean diameter considered is -7.65ϕ (i.e. $0.2\ m$) with a
 344 $\sigma_{Diam} = 1.2\phi$. In meters, the $\mu - \sigma$ and $\mu + \sigma$ are 0.09 and $0.46\ m$, respectively.
 345 The median ejection velocity was set to $100\ m\ s^{-1}$ with a $\sigma_{Vel} = 50\ m\ s^{-1}$.
 346 Although the median velocity is slightly lower than both the theoretical and
 347 observed velocities of Alatorre-Ibargüengoitia et al. (2012) for Popocatepetl
 348 volcano (Mexico), these authors suggest VBPs reaching maximum distances
 349 of $2.6\text{--}3.7\ km$. Based on the inversion modelling presented in Section 3.4
 350 and the reports of Mercalli and Silvestri (1891) and De Fiore (1922), VBPs
 351 reaching distances $\geq 1600\ m$ can be considered as outliers requiring ejection
 352 velocities comprised between 120 and $300\ m\ s^{-1}$ depending on the ejection
 353 angle. As a result, values of μ_{Vel} and σ_{Vel} proposed in Table 3 cover well
 354 these joint observations. The ejection angle was defined as a mean value
 355 centred on the vertical with a standard deviation of $\frac{\pi}{12}\ rad$, i.e. 15° .

356 A case of no wind was used here, and drag forces were accounted for
 357 using a standard atmosphere as described by Mastin (2001). A region of
 358 reduced drag was set with a radius of $200\ m$. Alatorre-Ibargüengoitia et al.
 359 (2012) report heights of about $600\ m$ at Popocatepetl volcano, which we chose
 360 to reduce since these explosions appear larger and characterized by higher
 361 ejection velocities and distances reached by VBPs. It is however important
 362 to notice that in the case of La Fossa, an altitude of $200\ m$ above the vent is

Table 3: Eruption source parameters associated with a Vulcanian-type eruption scenario at La Fossa volcano based on the eruption of 1888–1890. Different Gaussian distributions of densities are identified for ^a lithic blocks, ^b thin-rinded and ^c thick-rinded breadcrust bombs.

		Unit	Mean	σ
Source	Density	$kg\ m^{-3}$	2500	100^a
			800	50^b
			1600	200^c
	Diameter	ϕ	-7.65	1.2
	Velocity	$m\ s^{-1}$	100	50
	Ejection angle	rad	0	$\pi / 12$
	Number particles	—	10^6	—
Wind	Speed	$m\ s^{-1}$	0	—
	Direction	<i>Degrees</i>	0	—
Drag	Time step	s	0.01	—
	Pressure	hPa	1.01325×10^5	—
	Temperature at sea level	$^{\circ} K$	298	—
	Thermal lapse	$^{\circ} C\ km^{-1}$	-6.5×10^{-3}	—
	Reduced Drag radius	m	200	—

363 higher than the surrounding crater.

364 4.2. Probabilistic hazard assessment

365 The destructiveness caused by VBPs is mostly due to the associated high
366 kinetic energies at the impact, and the aim of this hazard assessment is
367 to investigate the probability to exceed critical energy thresholds. Various
368 thresholds, hereafter expressed as E_T (J), were estimated for the impact on
369 the built environment (e.g. Spence et al., 2005; Pomonis et al., 1999; Jenkins
370 et al., 2014) and will be discussed later. Since VBPs result in discontinuous
371 punctual impacts, it is necessary to average the number of impacts on a
372 representative area. However, since no standardized method yet exists, we
373 explore two different approaches for the quantification of the hazard related
374 to VBPs impacts.

375 4.2.1. Pixel-based approach

376 As a first approach, we average the VBP impacts on an equally-spaced
377 grid for each pixel of area $A_{i,j}$ in order to quantify the probability of occur-
378 rence a VBP of a given energy threshold in a given pixel:

$$P(A_{i,j}, E_T) = \frac{\sum VBP_{A_{i,j}, E_T}}{n_{VBP}}, \quad (10)$$

379 where n_{VBP} is the total number of simulated VBPs.

380 Since this approach introduces a dependency on the pixel area, we assess
381 the sensitivity of our post-processing method to i) the number of VBPs
382 simulated and ii) the resolution of the grid used to compile probabilistic
383 hazard assessments. The number of simulated VBPs was varied between 10^4
384 and 10^7 with increment steps of 10^1 . Grid resolutions of 5, 10, 20, 50, 75, 100,
385 200, 300, 400, 500, 600, 700, 800, 900 and 1000 m were tested. For each pair of
386 number of particles/grid resolution, 20 simulations were performed, resulting
387 in a total of 1,200 simulations. Probability calculations were performed for
388 three points, namely the top of the hiking path, the center of the Porto area
389 and Porto di Ponente, located at distances from the vent of 400, 1,300 and
390 1,700 m , respectively (green points 1, 2, and 3 on Fig. 1).

391 Figure 3 summarizes the sensitivity analysis performed using ESPs in Ta-
392 ble 3 and calculating the probability of VBPs exceeding an energy of 4000 J ,
393 identified as the minimum energy to penetrate weak RC slabs roofs (Spence
394 et al., 2005). Results show that:

- 395 • For a given point, an increase of the number of simulated particles does
396 not significantly affect the mean probability value but greatly reduces
397 the associated standard deviation (Fig. 3);
- 398 • For a given number of simulated particles, the probability decreases
399 with distance from the vent but the standard deviation remains in the
400 same order of magnitude;
- 401 • For the proximal point (i.e. Point 1 in Fig. 3), a change of order of
402 magnitude of mean probabilities (i.e. $10^{-2}\%$ to $10^{-1}\%$) occurs at a
403 resolution of about 200 m .

404 As a result, based on these observations and in order to find a good com-
405 promise between computation time and accuracy of the output, we simulate
406 10^6 particles averaged on a 100×100 m grid. In the absence of a plateau
407 with stable probability values, we fix the resolution threshold in the zone of
408 the lowest variability of mean probability values.

409 4.2.2. Zone-based approach

410 As a second approach, we assess the probability of impact in a zone of
411 interest Z . Here, such a zone is defined either as a distance from the vent (i.e.
412 the probability of impact at a given distance interval from the vent) or as a

413 radial sector (i.e. probability of impact at a given azimuth interval from the
414 vent). Probabilities of a VBP exceeding an energy threshold E_T can then be
415 expressed as normalized either on the total number of VBPs simulated or on
416 the number of VBPs that fell in a given zone Z . In the first case, $P(Z, E_T)$
417 answers the question "what is the probability of a VPB to exceed a given
418 energy threshold E_T in a zone Z ?". In the second case, $P(E_T|Z)$ answers the
419 question "knowing that a VBP impacts the zone Z , what is its probability
420 to exceed an energy threshold E_T ?".

421 Note that although the combination of both approaches might result in
422 an overall picture of the VBP hazard around a given volcano, the comparison
423 of the hazard with other volcanoes is difficult due to the nature of both the
424 modelling and the post-processing methods. Additionally, each approach to
425 the probabilistic quantification of the VBP hazard have different purposes.
426 For instance, the zone-based approach is more suitable for hazard zoning
427 purposes, whereas the pixel-based approach is more appropriate for impact
428 assessment purposes. For this reason, this latter one will be discussed in more
429 details in this paper, but the zone-based approach is thoroughly presented
430 in the user-manual of the GBF model.

431 *4.3. Vulnerability of the built environment*

432 VBPs can affect the built environment due to the dynamic impact caused
433 by their high kinetic energies, which can result in either damages to the struc-
434 tures, perforation or collapse of the building structure (Pomonis et al., 1999;
435 Blong, 1984; Jenkins et al., 2014; Spence et al., 2005). The propensity of
436 a building to suffer damages is typically expressed through fragility curves,
437 which describe the relationship between the intensity of the hazard and the
438 probability of damage. Such curves must be defined for each hazard and
439 each type of buildings. In the case of tephra fallout, the hazard is expressed
440 as the load caused by fallout leading to failure. For VBPs, the parameter
441 of importance is the kinetic energy at the impact sufficient for roof perfora-
442 tion. Fragility curves are usually defined as a combination of i) empirical
443 compilations of post-eruption observations of damages (e.g. Blong, 2003b;
444 Wilson et al., 2011; Pomonis et al., 1999), ii) laboratory experiments and
445 iii) theoretical studies on material strengths (e.g. Petrazzuoli and Zuccaro,
446 2004)

447 Here, we assess the vulnerability of the built environment to roof perfora-
448 tion from VBP impacts. The starting point of this study is the vulnerability
449 curves proposed by Spence et al. (2005) for the European area. Vulnerability

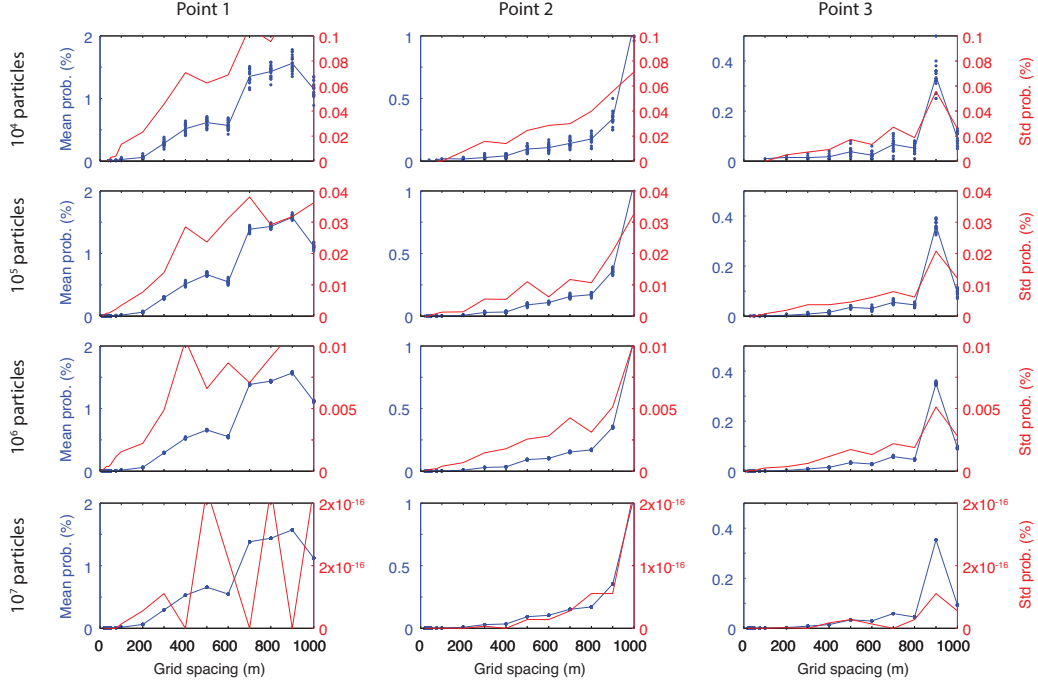


Figure 3: Sensitivity of the probabilistic hazard assessment strategy to i) the number of simulated particles and ii) the resolution of the grid used to quantify the probability of VBPs exceeding a given energy threshold E_T of 4000 J . Across sub-plots, the rows represent variable number of simulated particles and the columns represent the different points on which probabilities were calculated and include the top of the hiking path (Point 1), the center of the Porto area (Point 2) and Porto di Ponente (Point 3). Each plot has two y-axes: the left one (blue) shows the mean probability calculated over the 20 simulations (blue dots) for each set of number of particles/grid resolution; the right one (red) shows the corresponding standard deviation.

450 curves take the shape of a cumulative density function of a normal distribu-
 451 tion (ϕ) and are expressed as a function of the mean kinetic energy E_{mean}
 452 and a standard deviation σ assumed to have a fixed value of 0.2 (Spence
 453 et al., 2005). Following Spence et al. (2005) and Jenkins et al. (2014), the
 454 probability of perforation ($P_{perforation}$) is expressed as a function of the VBP
 455 energy I (J) with the following relationship:

$$P(Perforation|I) = \phi(\ln(I), \ln(E_{mean}), \sigma) \quad (11)$$

456 Biass et al. (submitted) provide a review of the built environment in
 457 Vulcano. The 2000 census of the Italian Instituto Nazionale di Statistica

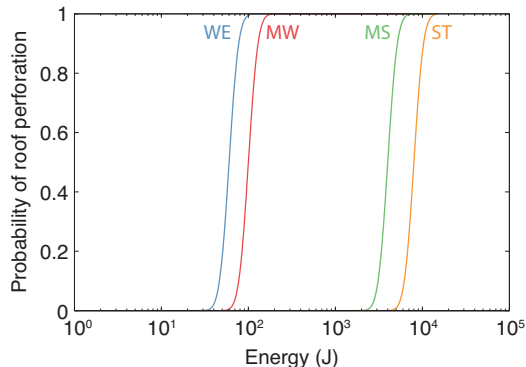


Figure 4: Fragility curves for the roof types *WE*, *MW*, *MS* and *ST* of Spence et al. (2005) as defined in Table 4.

458 (ISTAT, 2005) identifies 1093 buildings on the island, comprising 895 resi-
 459 dential houses and 64 public and tourism facilities. According to this census,
 460 the main construction period spans from the 1970's to 1980's, but discussions
 461 with inhabitants and workers on the island suggest that most buildings were
 462 renewed over the years, making the true period of construction difficult to
 463 assess. Additionally, the field survey realized in the context of the EU-funded
 464 *ENSURE* project (Bonadonna et al., 2011) provides detailed descriptions of
 465 the most representative building in a 100×100 m pixel, revealing that build-
 466 ing morphologies are homogeneously distributed over the settled areas and
 467 include 70% single-storey buildings, 73% with flat roofs and 54% with a reg-
 468 ular morphology. We assume that the majority of flat roofs are composite
 469 slabs made of precast reinforced concrete and clay bricks connected by an up-
 470 per layer of concrete and designed for access, suggesting a resistance to static
 471 loads of at least 150 kg m^{-2} (Jenkins et al., 2014). Additionally, our database
 472 on the built environment comprises the mapping building's footprints from
 473 aerial images (Galderisi et al., 2013).

474 Spence et al. (2005) identified three typical roofing stocks and their asso-
 475 ciated probabilities of occurrence over Europe and include *Weak*, *Median* and
 476 *Strong* roofing stocks, corresponding to exceedance probabilities of 10%, 50%
 477 and 90%, respectively. Each roofing stock is characterized by proportions of
 478 four roof types, namely *weak* (WE), *medium weak* (MW), *medium strong*
 479 (MS) and *strong* (ST), each defined by a specific fragility curve. Here, this
 480 method is extrapolated for i) the specific case of Vulcano and ii) to consider

Table 4: Description of the typical roofing stocks of Spence et al. (2005) adapted to the built environment of Vulcano (adjusted from Biass et al. (submitted)). The vulnerability of each roof class is characterized by a mean kinetic energy E_{mean} and a standard deviation σ fixed to 0.2. The E_{mean} is identified based on existing literature Spence et al. (2005); Tsunematsu (2012); Pomonis et al. (1999); Jenkins et al. (2014); Blong (1984); Mavrouli and Corominas (2010b). The proportions of each roof class in the composite European roofing stocks of Spence et al. (2005) are also shown assuming a median roofing stock (50% probability of occurrence) and a strong roofing stock (90% probability of occurrence). *RC stands for reinforced concrete.*

Roof class	Description	$E_{mean}(J)$	Proportion per roofing stock (%)	
			Median	Strong
WE (weak)	Tiled roof, poor condition	60	34.3%	2.7%
MW (medium weak)	Tiled roof, average or good condition	100	44.1%	18.9%
MS (medium strong)	Flat RC roof , average condition	4000	18.9%	44.1%
ST (strong)	Flat RC roof, good condition	8000	2.7%	34.3%

481 the impact related to VBPs. Firstly, as described by Biass et al. (submitted)
482 based on the different datasets described above, we assume that the popula-
483 tion of buildings in Vulcano are comprised between the typical *Median* and
484 *Strong* roofing stocks; the *Weak* stock is not considered here. Secondly, we
485 assume that buildings either have flat reinforced concrete roofs or tiled roofs
486 over a timber structure, which reflects well the field survey. Note that the
487 majority of buildings described by Biass et al. (submitted) consists of a flat
488 reinforced concrete roof with an additional tiled porch. Finally, the study
489 of Spence et al. (2005) focusing primarily on the impact related to tephra
490 fallout, fragility curves for each typical roof type (i.e. WE, MW, MS or
491 ST) are defined based on critical thresholds of static loads, which we need
492 to extrapolate here to express the vulnerability of the built environment to
493 VBPs.

494 Table 4 summarizes the roof classification adapted for the built environ-
495 ment on Vulcano, where the quality of both types of observed roofs (i.e.
496 reinforced concrete slabs and tiles) is either in average or good condition.
497 Following equation 11, we translate the approach usually taken to assess the
498 vulnerability to tephra fall to VBPs by defining fragility curves as a cumu-
499 lative function of a Normal distribution and characterized by a mean energy
500 $E_{mean} (J)$. Here, E_{mean} is defined as the ranges published in the existing
501 literature (e.g. Spence et al., 2005; Tsunematsu, 2012; Pomonis et al., 1999;
502 Jenkins et al., 2014; Blong, 1984) for the built environment approximat-
503 ing best the observations on Vulcano. Following the approach applied to

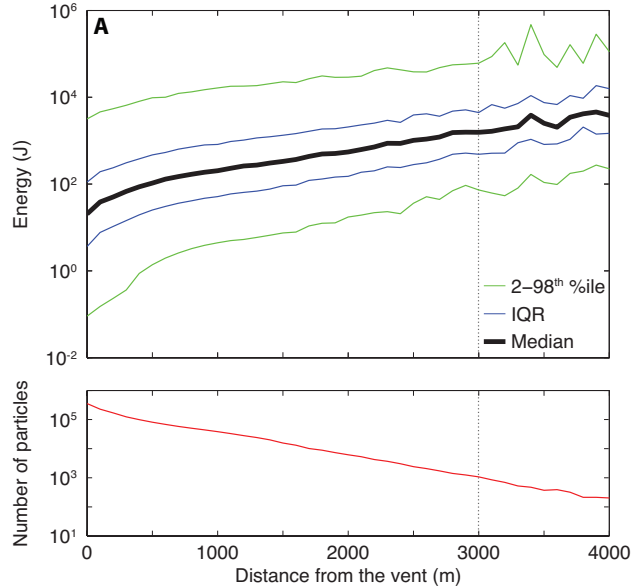


Figure 5: A: Median energy with distance from the vent. The uncertainty is expressed by the 25th–75th percentiles (i.e. IQR) and the 2nd–98th percentiles. The vertical dashed line indicates the distance (i.e. $\sim 3,000$ m) at which the number of particles is too limited to produce stable results. B: Number of particles with distance from the vent. A distance of $\sim 3,000$ m corresponds to $\sim 10^3$ particles.

504 tephra fallout, the standard deviation of the distribution (E_{std}) is fixed to
 505 0.2 (Spence et al., 2005; Jenkins et al., 2014).

506 5. Results

507 For the scenario ran with ESPs defined in Table 3, Figure 5A shows
 508 the variation of the median VBP energy with distance from the vent, with
 509 the associated variability expressed as the 25th–75th percentiles (i.e. IQR)
 510 and the 2nd–98th percentiles. Two main observations must be outlined from
 511 Figure 5A. Firstly, the median energy increases with distance from the vent,
 512 which is a consequence of the caprock assumption used to model Vulcanian
 513 explosions (Self et al., 1979; Wilson, 1980; Fagents and Wilson, 1993). Such
 514 an assumption implies that once the coherent plug reaches its fragmentation
 515 level (here considered as the reduced drag radius in Table 3), all VBPs are
 516 released with the same ejection velocity, regardless of their masses. As a
 517 result, only large VBPs possess a sufficient kinetic energy to reach distances

518 further away from the vent and are therefore associated with relative high
519 impact energies. Secondly, curves in Figure 5A follow a smooth trend up to
520 a distance of $\sim 3,000\text{ m}$ (i.e. vertical dashed line in Fig. 5), after which
521 the signal becomes chaotic. By projecting this distance on Figure 5B, it
522 can be observed that only 10^3 particles are falling at distances larger than
523 $\sim 3,000\text{ m}$ (i.e. 0.1% of the total number of simulated VBPs), which are not
524 sufficient to obtain stable results. Probabilities calculated for distances from
525 the vent larger than $\sim 3,000\text{ m}$ should thus be critically used.

526 5.1. Hazard assessment

527 Following the approach presented in section 4.2, probabilistic maps of
528 VBPs that exceed given hazardous energy thresholds of roof perforation were
529 first compiled. Figure 6A–B shows the geographical distribution of proba-
530 bilities to exceed impacts of 100 J (i.e. threshold for the perforation of tiled
531 roofs in average or good condition) and 4000 J (i.e. threshold for the perfo-
532 ration of reinforce concrete roofs in average condition). Impacts are averaged
533 on a $100 \times 100\text{ m}$ pixel and normalized over the total number of simulated
534 VBPs. Figure 6A–B shows an exponential decrease of the probability values
535 with distance from the vent. Secondly, the approach described in section
536 4.2.1 was applied to estimate probabilities of impact at a given distance from
537 the vent (Fig. 7A–B) or at a given radial sector around the vent (Fig. 7C–
538 D). Probabilities are expressed either as normalized over the total number of
539 simulated VBPs (i.e. $P(Z, E_T)$; Fig. 7A,C) or as normalized over the num-
540 ber of VBPs that impacted the considered zone (i.e. $P(E_T|Z)$; Fig. 7B,D).
541 Finally, hazard curves were compiled (Fig. 8), which show the probability of
542 exceeding any impact energy for the settled areas of Porto, Il Piano, Lentia
543 and Vulcanello (white squares in Fig. 1 and Fig. 6), located respectively 1.3,
544 2.4, 1.8 and 2.6 km away from the vent.

545 Results show that Porto, located 1.3 km North of the vent (Fig 1), is the
546 most exposed settlement with probabilities of impact on tiled and reinforced
547 concrete roofs of $8\text{--}9 \times 10^{-3}\%$ and $1\text{--}2 \times 10^{-3}\%$, respectively. The second most
548 exposed settlement is Lentia (1.8 km NE of vent), with probabilities of impact
549 $\sim 3 \times 10^{-3}\%$ (i.e. tiled roofs) and $\sim 5 \times 10^{-4}\%$ (i.e. reinforced concrete roofs).
550 The settlements of Il Piano and Vulcanello, located at respectively 2.4 km
551 SW and 2.6 km N of the vent (Fig 1), result in probabilities of $7 \times 10^{-4}\%$ and
552 $2 \times 10^{-4}\%$ of impact on tiled roofs, respectively, and negligible probabilities
553 of impact on reinforced concrete roofs. Note that all these probabilities are
554 calculated for a single geographic point in the center of all these settlement

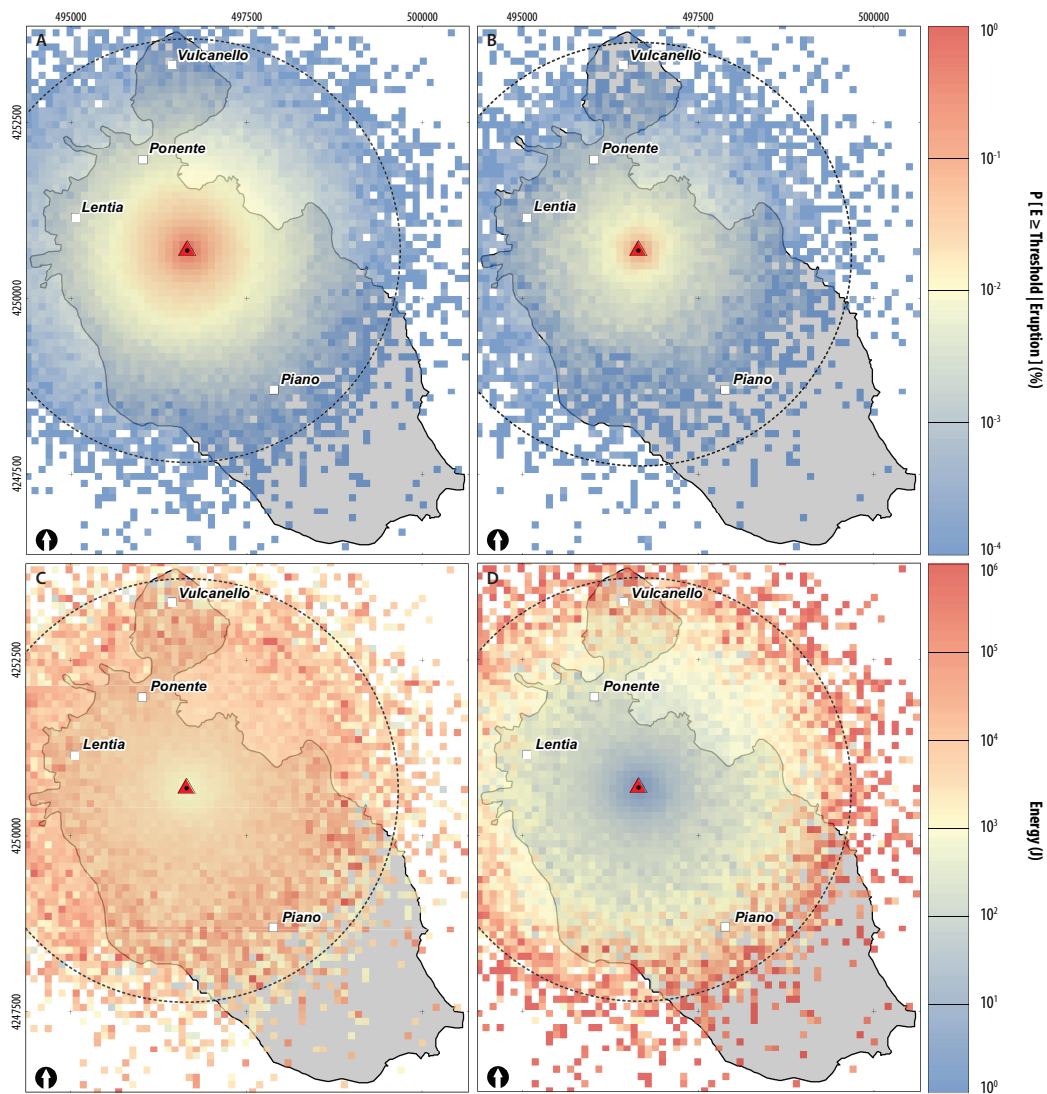


Figure 6: A–B: Probability maps (%) of VBPs exceeding energies of (A) 100 J and (B) 4000 J . D–E: Energies for probabilities of occurrence within a given pixel of (D) 10% and (F) 90%. The towns main are shown as white squares (Fig. 1). The black dashed line contours a distance of 3,000 m around the vent, considered as the distance beyond which not enough particles are observed to provide stable results (Fig. 5).

555 and are equal in the 100×100 m pixel surrounding the respective reference
 556 points.

557 When the zone-based approach is adopted to assess the probability of

558 impact at a given distance from the vent (Fig. 7A–B), the choice of the type
559 of probability (i.e. $P(Z, E_T)$ versus $P(E_T|Z)$) greatly influences the message
560 carried by the probabilistic hazard assessment. When normalized over the
561 total number of simulated VBPs, Figure 7A shows greater probabilities of
562 being impacted by a VPB with a kinetic energy of 4000 J close to the vent,
563 where a probability of 0.5% exists up to a distance of 1 km away from the
564 vent. In contrast, Figure 7B shows that should a VBP impact a given zone
565 of interest, there is a larger probability that it will exceed a kinetic energy
566 of 4000 J at larger distances from the vent. As a result, there is a 25%
567 probability that a VBP reaching a distance of 3,000 m from the vent will
568 exceed such an energy threshold. When a similar approach is applied on zones
569 of interest defined as radial sectors around the vent, no notable difference is
570 observed when using either type of probabilities, where $P(Z, 4000 J) =$
571 $0.2 - 0.5\%$ and $P(4000 J|Z) = 3 - 6\%$ all around the vent. This is possibly
572 due to the absence of any large topographic barrier around the La Fossa
573 cone.

574 In order to estimate the typical range of VBP energies at a given pixel,
575 maps showing the energy for a given probability of occurrence were also
576 produced (Fig. 6C–D). At each pixel, the 10th, 25th, 50th, 75th and 90th
577 percentiles were calculated over the energy of all VBPs that fell in a given
578 $100 \times 100 m$ area. Since the n^{th} percentile returns the lowest $n\%$ of the
579 population, there is a $100 - n\%$ probability that the energy will exceed the
580 energy given by the n^{th} percentile. As an illustration, the 10th percentile of
581 a given pixel shows the energy occurring with a 90% probability within this
582 given pixel. Note that this energy is based upon the conditional probability
583 that a VBP impact is occurring inside this pixel, and does not consider
584 the probability of the pixel to be impacted. Figure 6C–D illustrates the
585 geographical distributions of energies for probabilities of occurrence of 10%
586 and 90%, and reflect the increase of the average energy with distance from the
587 vent observed in Figure 5. Typically, when compared to large probabilities
588 of occurrence, lower probabilities of occurrence result in higher energies close
589 to the vent and lower energies further away. For instance, the crater area
590 is characterized by typical energies of $10^3 J$ and $10^1 J$ for probabilities of
591 occurrence of 10% and 90%, respectively, by opposition to $10^4 J$ and $10^3 J$
592 for the Porto area.

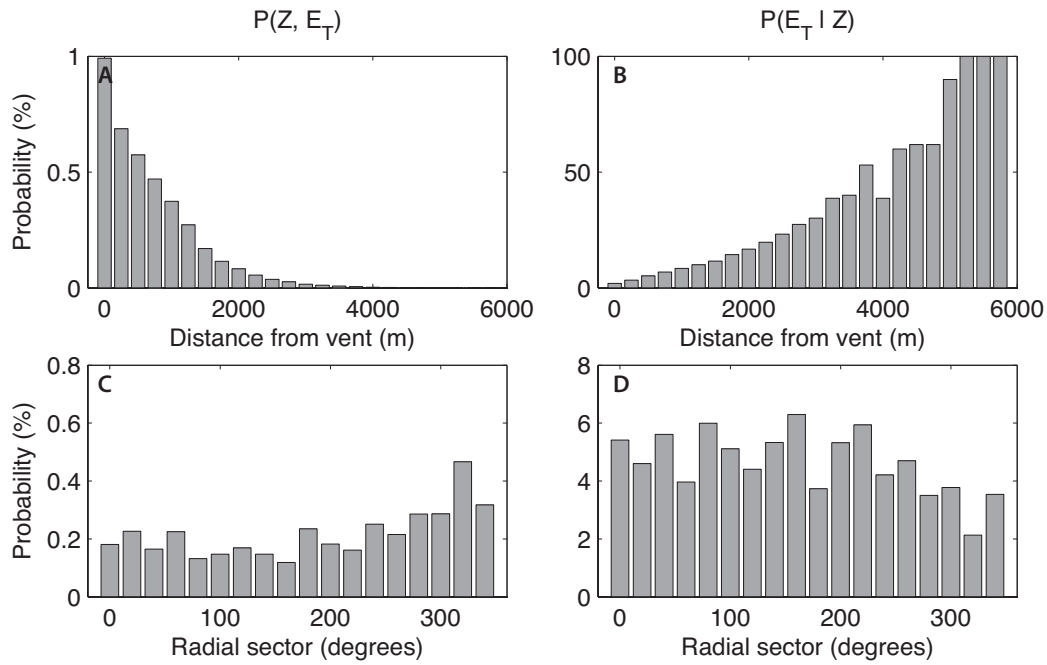


Figure 7: Probabilities of VBPs to exceed a kinetic energy of 4000 J at A–B a given distance from the vent and C–D a given radial sector around the vent. The left column expresses the probability normalized over the total number of simulated VBPs (i.e. $P(Z, E_T)$) whereas the right column is averaged over the number of VBPs that impacted the considered zone (i.e. $P(E_T | Z)$). Bins sizes are 250 m for the distance and 20° for the radial sectors.

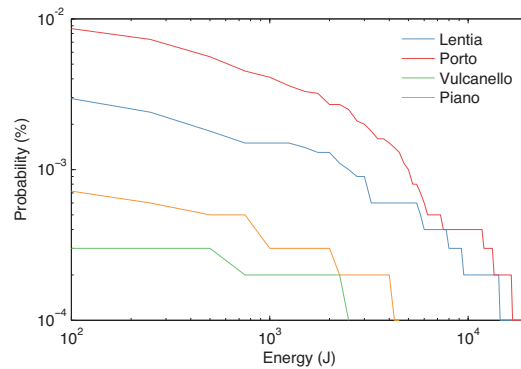


Figure 8: Hazard curves for the urban settlements areas shown as white squares in Fig. 1 and Fig. 6.

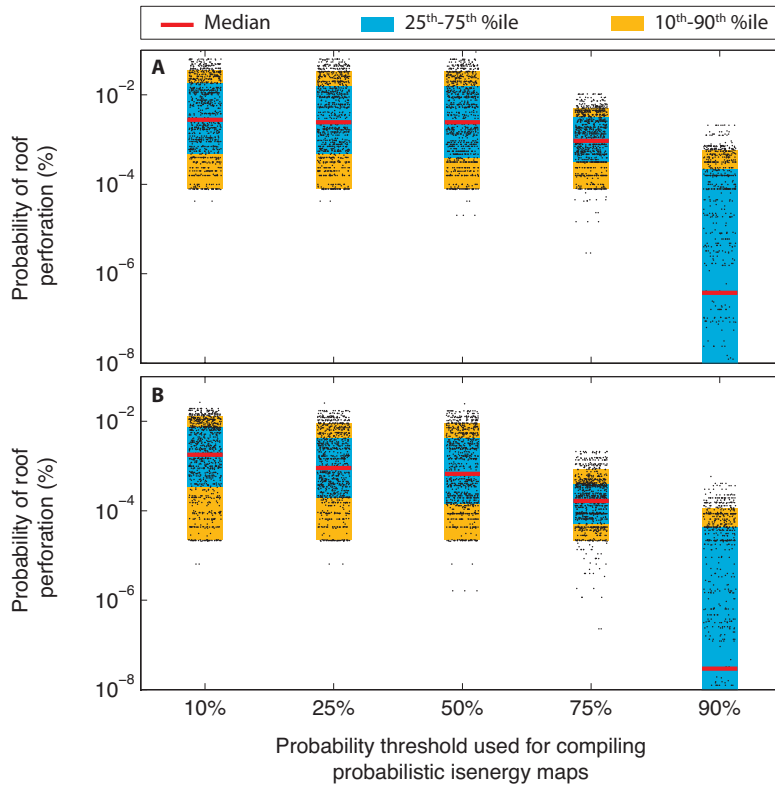


Figure 9: Impact on the built environment considering typical (A) median and (B) strong roofing stocks (Spence et al., 2005). The x-axis shows the probability threshold used to compile probabilistic energy maps. Black dots show the probability of roof collapse of each building resulting from the hazard occurring at a given probability (x axis). Distributions of probabilities of over all buildings are summarized as the median (red line), the 25th–75th interval (blue box) and the 10th–90th interval (orange box). For visibility, the lower y axis was manually set to 10⁻⁸%.

593 *5.2. Pre-event impact assessment*

594 The impact assessment was performed by combining the vulnerability
 595 curves shown in Table 4 and Figure 4 with the probabilistic energy maps (Fig.
 596 6C–D). For each building, energy thresholds of the containing pixel were
 597 retrieved for all probabilities of occurrence (i.e. 10%, 25%, 50%, 75% and
 598 90%). For each probability threshold, probabilities of roof perforation were
 599 calculated for a single building using the four curves defined by Spence et al.
 600 (2005) for the *WE*, *MW*, *MS* and *ST* roof classes (Table 4) and multiplied
 601 by the probability of occurrence of VBPs impact within the pixel of interest

Table 5: Summary of the impact considering a typical Median roofing stock (Spence et al., 2005). Values represent the probability of roof collapse calculated at given percentiles on the distributions shown in Fig. 9. P_{thresh} represents the probability threshold used to compile probabilistic energy maps (Fig. 6C–D).

Probability of roof perforation (%) – Median roofing stock					
Percentile:	10 th	25 th	50 th	75 th	90 th
$P_{thresh} = 10\%$	7.9×10^{-5}	4.9×10^{-4}	2.8×10^{-3}	1.8×10^{-2}	3.6×10^{-2}
$P_{thresh} = 25\%$	7.9×10^{-5}	4.7×10^{-4}	2.4×10^{-3}	1.5×10^{-2}	3.3×10^{-2}
$P_{thresh} = 50\%$	7.9×10^{-5}	3.9×10^{-4}	2.4×10^{-3}	1.5×10^{-2}	3.3×10^{-2}
$P_{thresh} = 75\%$	7.8×10^{-5}	3.1×10^{-4}	9.4×10^{-4}	3.1×10^{-3}	5.0×10^{-3}
$P_{thresh} = 90\%$	3.5×10^{-17}	1.6×10^{-13}	3.7×10^{-7}	2.1×10^{-4}	5.9×10^{-4}

602 (i.e. $n_{VBP, pixel\ x,y} / n_{VBP\ total}$). Following the typical roofing stocks of Spence
603 et al. (2005) for the European area, the resulting probabilities were then
604 multiplied by the probabilities of occurrence of their respective roof classes
605 in either Median or Strong roofing stocks (Table 4) and summed into a final
606 composite probability of roof perforation.

607 Figure 9, Table 5 and Table 6 summarize the impact of VBPs on the
608 built environment. Figure 9 can be read as a box and whisker plot, in which
609 black dots indicate raw composite probabilities of perforation of individual
610 buildings ($n = 1093$) calculated for a hazard occurring at a given probability
611 of occurrence (x axis). The resulting distributions are displayed as the me-
612 dian (red line), the 25th–75th percentiles range (blue area) and the 10th–90th
613 percentiles range (orange area). For instance, considering i) a probability of
614 occurrence of the hazard of 50% and ii) a median roofing stock, the top of
615 the orange box in Figure 9 (i.e. the 90th percentile) shows that 90% of the
616 buildings have a $\leq 3.3 \times 10^{-2}\%$ probability of roof perforation, which also
617 means that the remaining 10% have a $> 3.3 \times 10^{-2}\%$ probability of perfora-
618 tion. Tables 5 and 6 report the same type of information. For i) a probability
619 of occurrence of the hazard of 50% and ii) a strong roofing stock, the 50th
620 percentile in Table 6 shows that 50% of the buildings have probabilities of
621 roof perforation either lower or higher than $6.7 \times 10^{-4}\%$. Results indicate
622 that probabilities of roof perforation are lower than 0.1% for all roofing stocks
623 and probabilities of occurrence of the hazard.

Table 6: Summary of the impact considering a typical Strong roofing stock (Spence et al., 2005). Values represent the probability of roof collapse calculated at given percentiles on the distributions shown in Fig. 9. P_{thresh} represents the probability threshold used to compile probabilistic energy maps (Fig. 6C–D).

Probability of roof perforation (%) – Strong roofing stock					
Percentile:	10^{th}	25^{th}	50^{th}	75^{th}	90^{th}
$P_{thresh} = 10\%$	2.2×10^{-5}	3.5×10^{-4}	1.8×10^{-3}	7.5×10^{-3}	1.3×10^{-2}
$P_{thresh} = 25\%$	2.2×10^{-5}	2.0×10^{-4}	9.0×10^{-4}	4.2×10^{-3}	9.2×10^{-3}
$P_{thresh} = 50\%$	2.2×10^{-5}	1.4×10^{-4}	6.7×10^{-4}	4.2×10^{-3}	9.1×10^{-3}
$P_{thresh} = 75\%$	2.1×10^{-5}	5.2×10^{-5}	1.7×10^{-4}	3.9×10^{-4}	8.4×10^{-4}
$P_{thresh} = 90\%$	2.8×10^{-18}	1.3×10^{-14}	2.9×10^{-8}	4.3×10^{-5}	1.1×10^{-4}

624 6. Discussion

625 We introduce a new model called *The Great Balls of Fire* designed for the
626 probabilistic analysis of VBP impacts. The model relies on the identification
627 of probabilistic eruption scenarios described by distributions of selected in-
628 put parameters, namely i) initial ejection velocities, ii) size distribution and
629 iii) densities of VBPs. Sets of post-processing functions are also provided
630 to compile probabilities of VBP impacts exceeding hazardous thresholds of
631 kinetic energies. Probabilities can be expressed on a pixel-based approach,
632 suitable for hazard and pre-event impact assessments, or on zones of interests
633 (either concentric circles around or radial sector around the vent), suitable
634 for hazard zoning purposes.

635 6.1. Probabilistic hazard assessment for VBPs

636 Hazard assessments for VBPs published in the literature follow two main
637 approaches. Some authors used the *Eject!* model to estimate probability
638 density functions of impact distances based on ESPs inferred from observed
639 VBPs (e.g. Sandri et al., 2014). In contrast, other authors associate hazard
640 zones based on deterministic eruption scenarios with their respective proba-
641 bilities of occurrence (e.g. Alatorre-Ibargüengoitia et al., 2006, 2012). Here,
642 we aim at providing a fully probabilistic assessment for VBP impacts as a ba-
643 sis to produce long-term multi-hazard assessments based on Bayesian event
644 trees (e.g. Marzocchi et al., 2008; Selva et al., 2010; Sandri et al., 2014; Shel-
645 drake, 2014; Thompson et al., 2015). The probabilistic approach adopted
646 here is associated with a dependency on both the number of simulated VBPs
647 and on the size of the zones of interest defined to average VBP impacts.

648 This aspect should be investigated on a case-per-case basis, with the aim of
649 finding the best compromise between computation time and output accuracy.
650 As an example at La Fossa, Figure 3 shows minimum discrepancies of mean
651 and standard deviation values of probabilities from 10^6 simulated particles,
652 which results in valid results up to a distance 3,000 m away from the vent,
653 shown as the dashed circle on Figure 6. In contrast, 10^7 particles increase
654 the confidence radius to about 3,500 m , but results in both calculation and
655 post-processing times multiplied by a factor 10.

656 *6.2. Probabilistic eruption scenarios for VBPs*

657 In probabilistic hazard assessments, eruption scenarios are typically ex-
658 pressed as distributions of the most critical ESPs for the modelled phe-
659 nomenon (e.g. earthquake source parameters for seismic and tsunami hazard
660 assessments, Geist and Parsons, 2006; volume for landslide hazard assess-
661 ments, Guzzetti et al., 2005; thickness and volumes for lava flows, Connor
662 et al., 2012). Alatorre-Ibargüengoitia et al. (2006) identified the total kinetic
663 energy of Vulcanian explosions as the relevant ESP for defining eruption sce-
664 narios for VBPs, which can practically only be relevant when i) the ballistic
665 model is coupled with a conduit model (e.g. Alatorre-Ibargüengoitia et al.,
666 2012) and ii) when deterministic eruption scenarios are used.

667 Eruption scenarios as defined with our method differ from those presented
668 by Alatorre-Ibargüengoitia et al. (2012) for Popocatepetl on two main points.
669 Firstly, in our method, ESPs are those identified by Mastin (2001) stochas-
670 tically sampled on a Gaussian distribution (Table 3). Secondly, the hazard
671 zones resulting from the hazard assessment of Alatorre-Ibargüengoitia et al.
672 (2012) for Popocatepetl are a direct consequence of the eruption scenarios,
673 and, for instance, the high-hazard zone is defined as the typical VBP range
674 resulting from the most likely and least intense type of activity. This de-
675 terministic approach, although complementary to the probabilistic approach
676 when the probability of a future eruption tends to 1 (Marzocchi et al., 2008),
677 is of limited information for long-term planning and risk reduction strategies.
678 As an example, the cone of Popocatepetl is mostly deserted within a radius
679 of a few kilometres around the vent, and the purpose of a risk assessment
680 for VBPs is mainly the delimitation of exclusion zones. In contrast, urban
681 areas are found within a radius of 1 km around La Fossa and probabilistic
682 approaches become a necessity to estimate the likelihood of occurrence of
683 VBPs impacts as a first step towards the development and implementation
684 of pro-active risk mitigation strategies.

685 *6.3. Eruptive scenarios at La Fossa*

686 We developed a scenario for typical long-lasting Vulcanian eruptions at
687 La Fossa based on the inversion of field observations (Fig. 2) and the com-
688 parison with published literature (e.g. Alatorre-Ibargüengoitia et al., 2012;
689 Fitzgerald et al., 2014; Tsunematsu et al., 2014). Using the caprock assump-
690 tion, VBPs of different sizes have equal probabilities to be launched in the
691 velocity range expressed in Table 3. Ejection velocities reported in the lit-
692 erature range from 30 to 400 $m s^{-1}$ (Mastin, 1995; Fitzgerald et al., 2014;
693 Alatorre-Ibargüengoitia et al., 2012). In the case of La Fossa, the distribution
694 was assumed Gaussian with values of mean and standard deviations of 100
695 and 50 $m s^{-1}$, respectively, which implies that 95% of the VBP's will result
696 in ejection velocities comprised between >0 and 200 $m s^{-1}$, respectively. We
697 argue that this range is justifiable because i) it covers the majority of ejection
698 velocities identified for other volcanoes while discarding sub- or supersonic
699 velocities that are unlikely at La Fossa and ii) agrees with ranges obtained
700 through inversion of field data (Fig. 2). The size distribution of VBPs is
701 described here by a Gaussian distribution in ϕ units (i.e. a log-normal dis-
702 tribution in metres). Although Weibull distributions or fractal dimensions are
703 thought to better represent the physical fragmentation process (Brown and
704 Wohletz, 1995; Kaminski and Jaupart, 1998; Turcotte, 1986), the assump-
705 tion of a Gaussian shape is frequent in numerical modelling (e.g. Macedonio
706 et al., 1988; Fitzgerald et al., 2014).

707 At La Fossa, the 1888–1890 eruption is characterized by at least three
708 populations of VBPs characterized by different densities (Table 2). Our ap-
709 proach accounts for three different populations of densities, weighing the
710 number of simulated VBP according to proportions of occurrence of each
711 VBP type observed in the field. However, observations of Mercalli and Sil-
712 vestri (1891) suggest that each VBP type was produced at different stages of
713 the two-year-long Vulcanian cycle. Outcomes of our probabilistic hazard as-
714 sessment do not capture the evolution of VBP type through time and should
715 be viewed as a time-integrated hazard over the duration of a Vulcanian cycle.

716 *6.4. VBP hazard for Vulcano*

717 At La Fossa, Figure 6A–B shows a maximum probability of occurrence
718 of 1% for an energy threshold of 100 J . The minimum probability of 10^{-4}
719 (Fig. 6) is constrained by the number of simulated particles and occurs when
720 a given pixel was impacted by one single VBP. Such low probabilities result
721 from the reasons discussed in section 6.1, i.e. the consideration of a hazard

722 occurring on discrete points rather than a continuous blanketing resulting,
723 for instance, from tephra fallout. In this latter case, a probability of 100% of
724 a threshold of tephra accumulation occurs when this threshold was exceeded
725 at a given pixel at each single run of the model. In contrast, an hypothet-
726 ical probability of 100% to exceed a given energy threshold when consider-
727 ing VBPs would imply that all simulated particles exceeded the probability
728 threshold and fell into a single pixel. As a result, although Biass et al. (sub-
729 mitted) show an average probability 15–30% to exceed critical accumulations
730 of tephra for the collapse of the weakest roofs in the Porto area, probabilities
731 of occurrences of VBPs with critical energies for the built environment are
732 comprised between $10^{-3} - 10^{-2}\%$. When probability maps are converted to
733 energy maps (Fig. 6), our results show a probability of occurrence of high
734 energies that increases with distance from the vent. For the case of a steam-
735 blast eruption, Dellino et al. (2011) suggest a zone extending 200 *m* from
736 the vent characterized by a maximum energy of 10^6 *J* based on field obser-
737 vations. Our probabilistic approach suggests that in the case of a Vulcanian
738 eruption, such an energy has low probabilities of occurrence within 2,500 *m*
739 around the vent.

740 La Fossa is a 391 *m* high cone surrounded in the south and west by a
741 caldera rim rising from 250 to 400 *m* above sea level. From the DEM, the
742 height of the actual crater was estimated at ~ 220 *m*, and GBF simulations
743 were performed with a 200 *m*-high region of reduced drag (Table 3). As a
744 result, although a careful observation of the distribution of probabilities on
745 Figure 6 reveals a slight increase of impacts on the caldera rim, our hazard
746 assessment shows that the island does not host significant topographic bar-
747 riers to shelter from VBPs, leaving only the southernmost part of the island
748 with a virtually null probability of impact. On the other hand, simulations
749 performed with a null region of reduced drag does not significantly change
750 the final probabilistic hazard assessment.

751 Biass et al. (submitted) presents a study of wind patterns for the period
752 1980–2010 inferred from the ECMWF ERA- Interim database (Dee et al.,
753 2011), which reveals a $\sim 70\%$ probability of wind directed towards SE at sea
754 level, with associated velocities rarely higher than 20 *m s*⁻¹. To test the
755 influence of wind on the final probabilistic hazard assessment, simulations
756 were run with a mean wind with a constant velocity of 20 *m s*⁻¹ and a
757 constant wind direction (i.e. provenance + 180°) of 135°. Results show that
758 the final probabilities are not significantly affected by wind conditions. This
759 is due on one side to the fact that smaller particles will be more influenced

760 by wind forces, which will necessarily fall relatively close to the vent due to
761 the caprock assumption. In this case the large number of particles falling in
762 proximal area is the dominant influence on the final probability values. On
763 the other side, only a limited number of large particles will impact more distal
764 areas, but since wind has little effect on them, their additional displacement
765 is not sufficient to affect the final probability values.

766 *6.5. Pre-event impact assessment*

767 The probabilistic hazard assessment was combined with an estimation of
768 the vulnerability of the built environment to assess potential impacts asso-
769 ciated with VBPs. Here, the vulnerability regards the potential strength of
770 a building to resist roof perforation resulting from a dynamic impact, which
771 regards the risk on the built environment as a potential loss of life (e.g.
772 Spence et al., 2005; Jenkins et al., 2014) rather than expressing the loss of
773 economical value (e.g. Blong, 2003a). Following Biass et al. (submitted),
774 the vulnerability of the built environment was based on the typical roofing
775 stocks of Spence et al. (2005) for the European area, extrapolated to dynamic
776 impacts following two main assumptions.

777 Firstly, the limited observations of damages related to VBPs impacts
778 does not allow to develop robust vulnerability curves. In natural hazards,
779 the closest analogous phenomena associated with impacts at high kinetic en-
780 ergies include hail storms and rockfalls (e.g. Andrews and Blong, 1997; Hohl
781 et al., 2002; Mavrouli and Corominas, 2010a,b; Agliardi et al., 2009). Re-
782 sulting fragility curves can take various shapes such as sigmoid (e.g. Agliardi
783 et al., 2009) and logistic (e.g. Hohl et al., 2002) shapes. Here, in the absence
784 of more detailed information, we follow the approach undertaken for tephra
785 fallout (e.g. Spence et al., 2005; Pomonis et al., 1999; Jenkins et al., 2014)
786 using a lognormal distribution and a fixed geometric standard deviation of
787 0.2. Secondly, published post-event impact assessments report VBP im-
788 pacts associated with variable energy thresholds (e.g. Pomonis et al., 1999;
789 Blong, 2003b, 1984). Here, we estimated mean energy thresholds for the
790 built environment on Vulcano by comparing observed impacts with typology
791 of buildings resulting from our field survey (Biass et al., submitted). As a
792 result, two end-members of vulnerability to VBPs were identified compris-
793 ing tile roofs on the weakest spectrum and reinforced concrete roofs on the
794 strongest. Figure 4 reflects this bipolarity due to critical energy thresholds
795 varying by orders of magnitude between the two families of roofs identified
796 in Vulcano (i.e. tiles and reinforced concrete; Table 4).

797 In terms of cascading effects between volcanic hazards, the relationship
798 between VBPs and tephra is ambiguous. On one hand, tephra can act as a
799 blanket absorbing energy from a VBP and thus reduce its propensity to per-
800 foration from a dynamic impact. On another hand, VBPs can increase the
801 static load already caused by tephra layers and contribute to roof collapse.
802 These complex vulnerability patterns occurring in the context of multi-
803 hazards risk assessments were already discussed by Zuccaro et al. (2008) and
804 underline the complex task of combining fragility curves for different natures
805 of hazards (i.e. static load vs. dynamic impact) potentially simultaneously
806 affecting exposed elements.

807 7. Conclusion

808 A new approach for the hazard assessment related to the ejection of VBPs
809 is introduced, which quantifies the probabilities of occurrence of VBP impacts
810 exceeding hazardous thresholds of kinetic energy. This approach, in line with
811 recent efforts to quantify volcanic hazards in terms of probabilities, relies on
812 a new ballistic model called *The Great Balls of Fire*, with the main features
813 being:

- 814 • The definition of ESPs in terms of probability distributions;
- 815 • A variable drag coefficient;
- 816 • A fast computation time;
- 817 • The possibility to work on single CPUs or clusters of computers;
- 818 • Platform independent.

819 The model is distributed in open-source and made available on *VHub*. It
820 was validated using field observations of VBPs associated from eruptions of
821 La Fossa volcano. Additionally, sets of *Matlab* functions are provided to post
822 process the model output into probabilistic hazard assessments for VBPs,
823 resulting in a format useful for the integration in various GIS environments.
824 Both the model and the post-processing functions are available as Online
825 Resources along with the user manual.

826 A generic Vulcanian eruption scenario was identified for La Fossa based
827 on the stratigraphy of the last 1000 years. Results show that the settlements
828 of Lentia and Porto are the most likely to be impacted by VBP, whereas

829 Vulcanello and Piano are relatively safer (Fig. 4). In addition, the vulner-
830 ability of the built environment was assessed by extrapolating the generic
831 vulnerability curves for Europe of Spence et al. (2005) for tephra fallout to
832 VBPs based on a review of critical energy thresholds found in the literature
833 along with a field survey of the built environment on Vulcano. Both hazard
834 and vulnerability aspects were then combined to produce a first-order pre-
835 event impact assessment in terms of potential number of affected buildings.
836 Results show that although no building has a probability of roof perforation
837 higher than $10^{-2}\%$, the occurrence of a VBP impact is likely to result in high
838 repair costs.

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